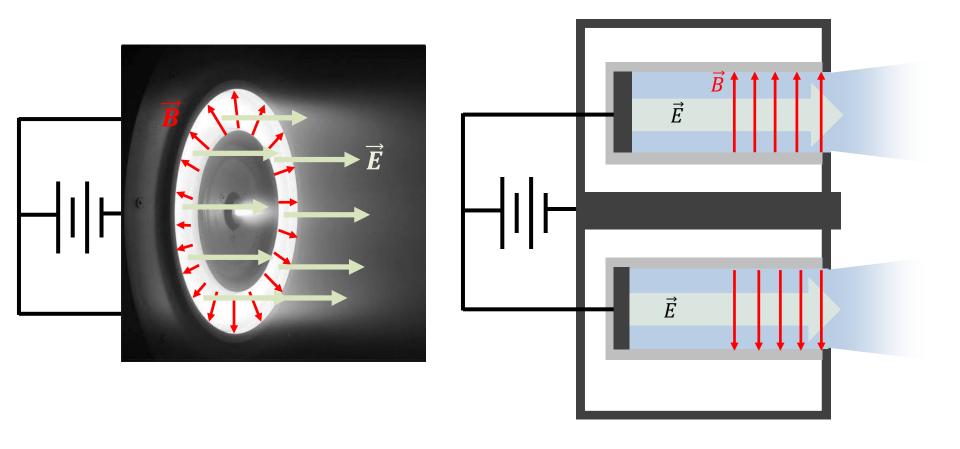


Probe-based Measurements of High-Frequency Azimuthal Oscillations in a Magnetically Shielded Hall Thrusters

Benjamin Jorns
Zachariah Brown
University of Michigan
Princeton University ExB Workshop



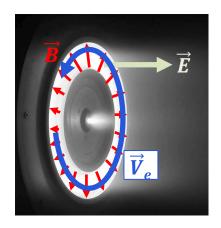
The Hall effect thruster



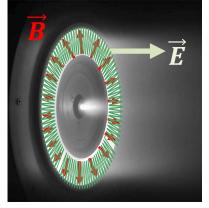
- What drives the anomalous across-field transport?
- **Driving hypothesis:** transport results from onset of microturbulence in $E \times B$ direction



Mechanism for how onset of microturbulence can drive cross-field transport

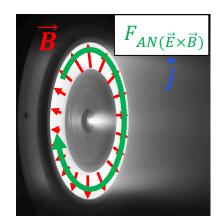


1) Strong E × B drift between electrons and ions

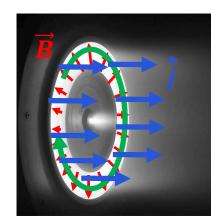


2) Azimuthal electron cyclotron drive instability (ECDI) driven unstable by drift through inverse cyclotron or Landau damping

3) Electrons slowed in E × B direction by wave growth leads to effective drag in Hall direction

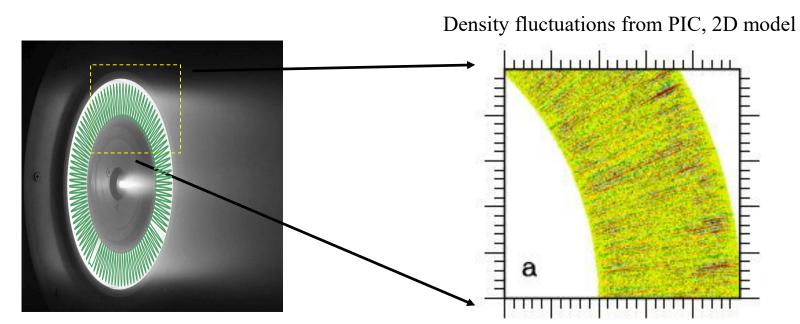


4) Effective drag due to onset of waves gives rise to cross-field electron current





Kinetic simulations predict onset of micro-turbulence



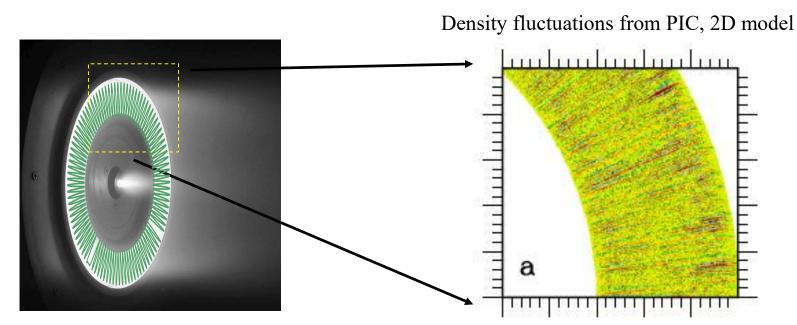
Sampling of other numerical models showing instability

- J. C. Adam , A. Héron , and G. Laval, Physics of Plasmas 11 , 295 (2004)
- A. Ducrocq, J. C. Adam, A. Héron, and G. Laval, Physics of Plasmas 13, 102111 (2006);
- J.P. Boeuf. Frontiers in Physics, Vol. 2, No. 74, (2014)
- T. Lafleur, , S. D. Baalrud, and P. Chabert, Physics of Plasmas 23, 053502 (2016);
- V Croes et al. Plasma Sources Sci. Technol. 26 (2017)
- Janhunen, S., et al., Physics of Plasmas 25, (2018)

* A. Héron and J. C. Adam, "Anomalous conductivity in Hall thrusters: Effects of the non-linear coupling of the electron-cyclotron drift instability with secondary electron emission of the walls. "Physics of Plasmas 20, 082313 (2013);



Kinetic simulations predict onset of micro-turbulence



Sampling of other numerical models showing instability

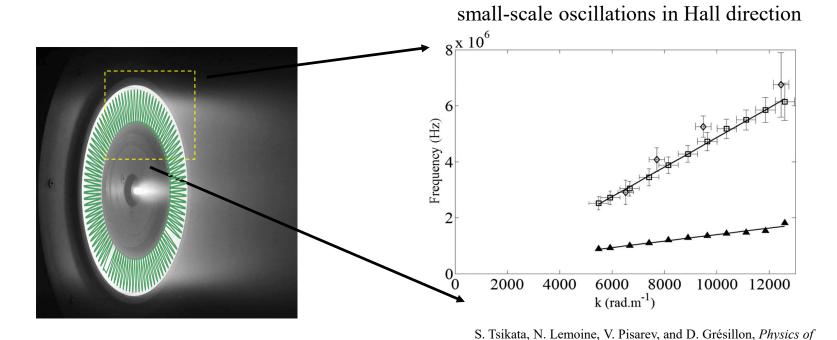
- J. C. Adam , A. Héron , and G. Laval, Physics of Plasmas 11 , 295 (2004)
- A. Ducrocq, J. C. Adam, A. Héron, and G. Laval, Physics of Plasmas 13, 102111 (2006);
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- V Croes et al. Plasma Sources Sci. Technol. 26 (2017)
- S. Janhunen et al., *Physics of Plasmas*, 011608 (2018)

* A. Héron and J. C. Adam, "Anomalous conductivity in Hall thrusters: Effects of the non-linear coupling of the electron-cyclotron drift instability with secondary electron emission of the walls. "Physics of Plasmas 20, 082313 (2013);

Does this type of instability actually exist in Hall thruster discharges?



Experimental evidence of microturbulence



- Wavelengths < 1 mm
- Dispersion is acoustic-like
- Modes are incoherent

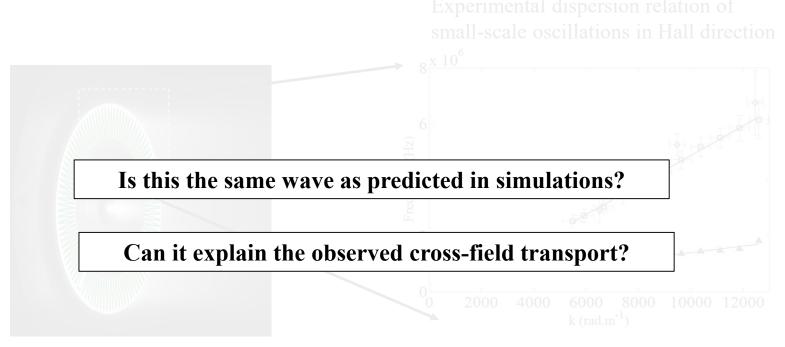
ECDI in the acoustic-like limit

Plasmas. Vol. 16., No. 3. 2009.

Experimental dispersion relation of



Experimental evidence of microturbulence



S. Tsikata, N. Lemoine, V. Pisarev, and D. Grésillon, *Physics of Plasmas*. Vol. 16., No. 3. 2009.

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ECDI in the acoustic-like limit

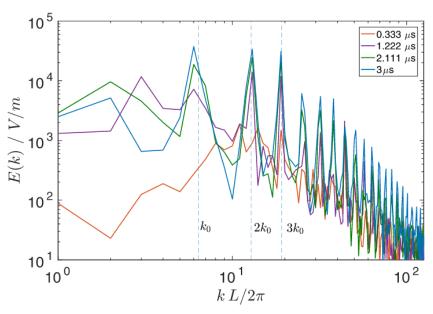
Modes are incoherent



Open experimental questions related to ECDI in Hall thrusters

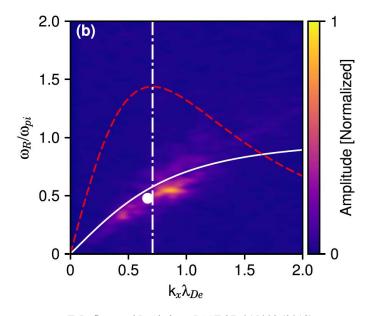
Is this the same wave as predicted in experiments:

What is the wavelength/frequency of maximum growth?



S. Janhunen et al., Physics of Plasmas, 011608 (2018)

Max growth at cyclotron resonance



T. Lafleur and P. Chabert, PSST 27 015003 (2018)

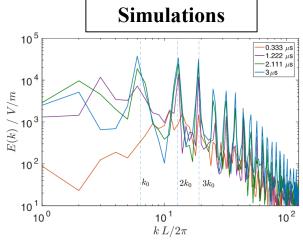
Max growth on order of Debye length



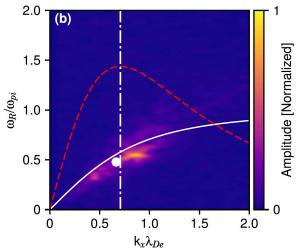
Open experimental questions related to ECDI in Hall thrusters

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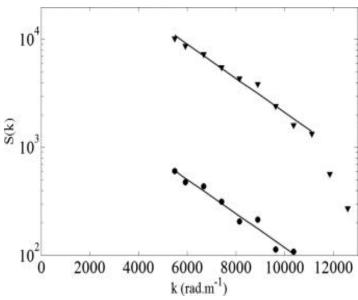


S. Janhunen et al., Physics of Plasmas, 011608 (2018)



T. Lafleur and P. Chabert, PSST 27 015003 (2018)

Experimental measurements



S. Tsikata, N. Lemoine, V. Pisarev, and D. Grésillon, *Physics of Plasmas*. Vol. 16., No. 3. 2009.

Contradictions with simulations

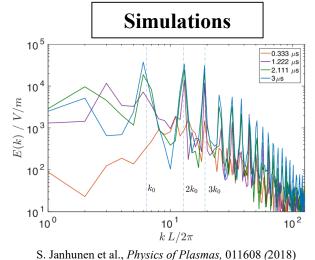
- Measurements are broadband
- Do not show maximum growth wavelength

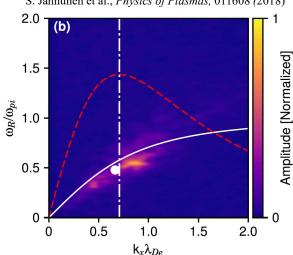


Open experimental questions related to ECDI in Hall thrusters

Is this the same wave as predicted in experiments:

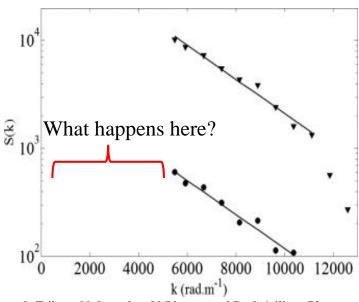
What is the wavelength/frequency of maximum growth?





T. Lafleur and P. Chabert, PSST 27 015003 (2018)

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S. Tsikata, N. Lemoine, V. Pisarev, and D. Grésillon, *Physics of Plasmas*. Vol. 16., No. 3. 2009.

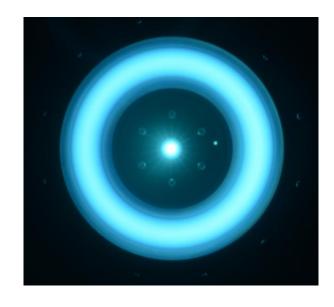
Contradictions with simulations

- Measurements are broadband
- Do not show maximum growth wavelength



Experimental Setup at UM

H9: 9kW magnetically shielded Hall effect thruster at 300 V and 4.5 kW



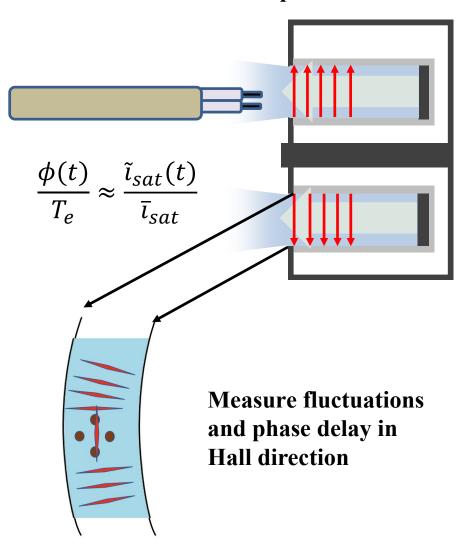
Large Vacuum Test Facility at UM





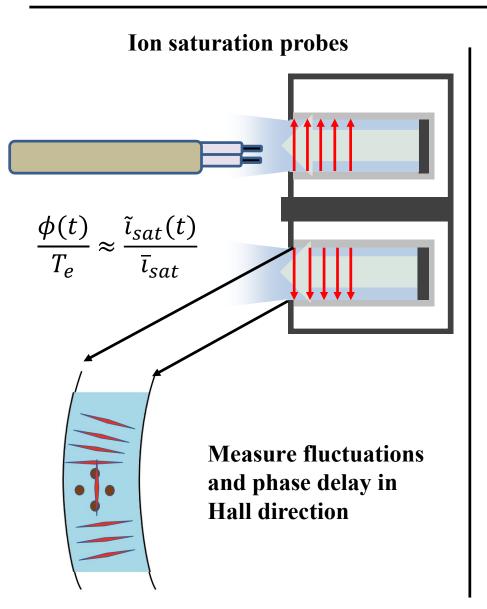
Experimental Setup at UM

Ion saturation probes



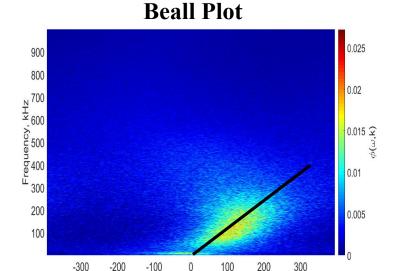


Experimental Setup at UM



Analysis

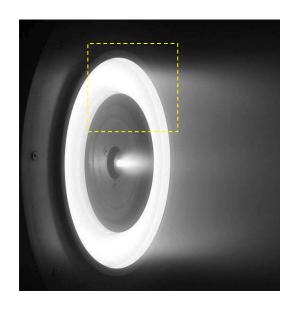
- Fourier analysis to find
 - Power spectrum $\phi(\omega)$
 - Cross-correlation $\omega(k)$
- Averaging yields statistical dispersion relation $\phi(\omega, k)$

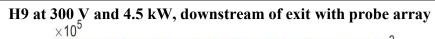


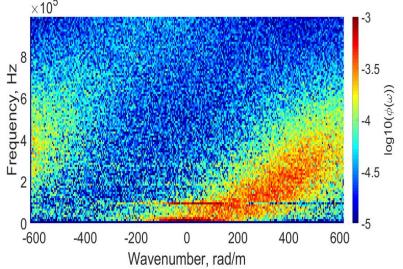
Z. Brown and B. Jorns, AIAA-2018-4423

Wavenumber, m⁻¹

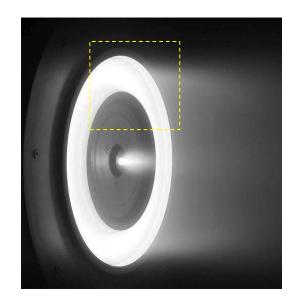




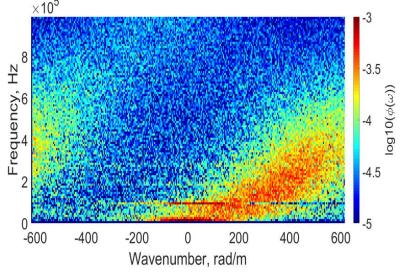




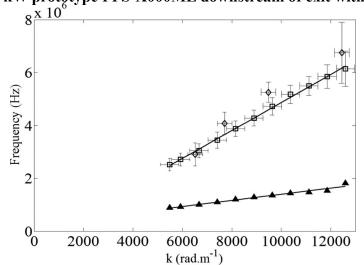




H9 at 300 V and 4.5 kW, downstream of exit with probe array

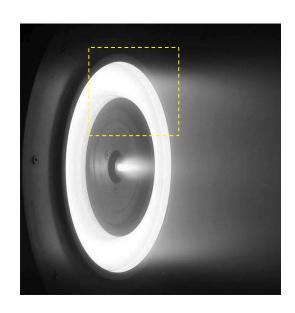


5 kW prototype PPS-X000ML downstream of exit with CTS

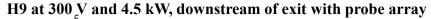


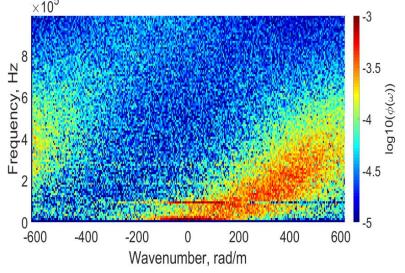
S. Tsikata, N. Lemoine, V. Pisarev, and D. Grésillon, *Physics of Plasmas*. Vol. 16., No. 3. 2009.



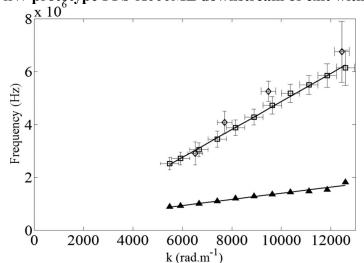


- Linear dispersion confirmed
- Results fill in some missing wavenumber space



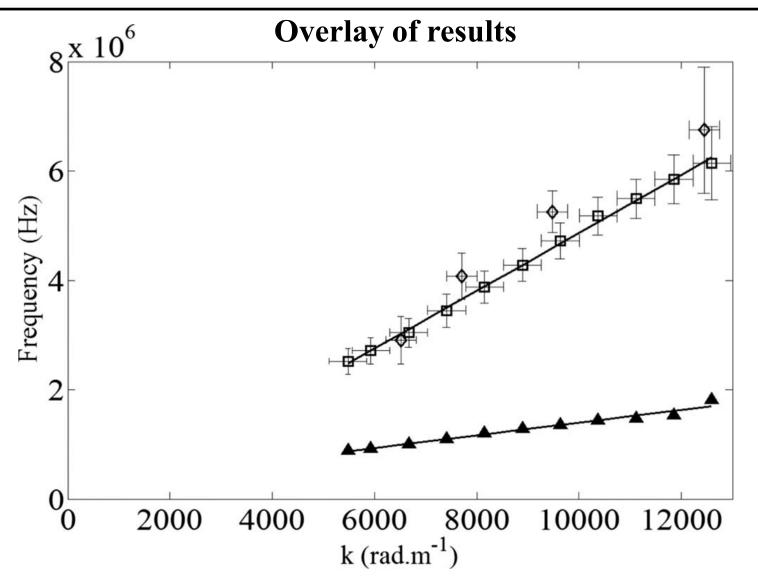


5 kW prototype PPS-X000ML downstream of exit with CTS

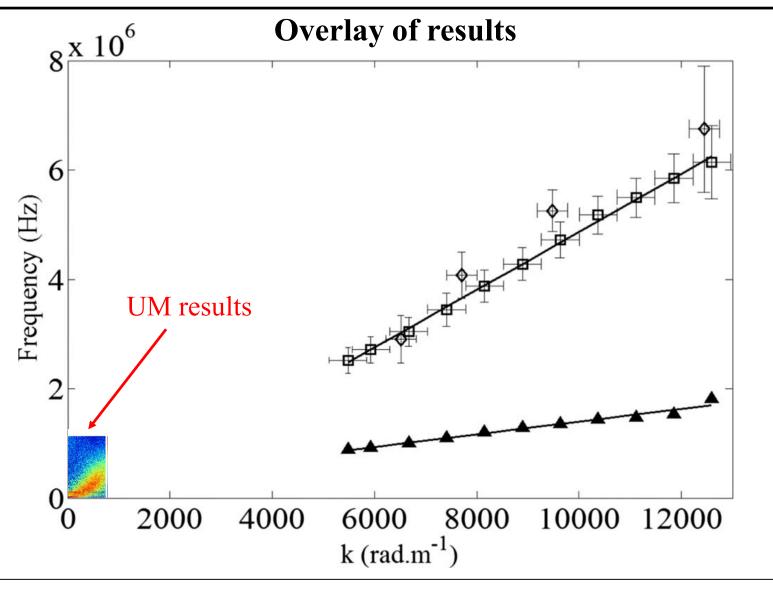


S. Tsikata, N. Lemoine, V. Pisarev, and D. Grésillon, *Physics of Plasmas*. Vol. 16., No. 3. 2009.



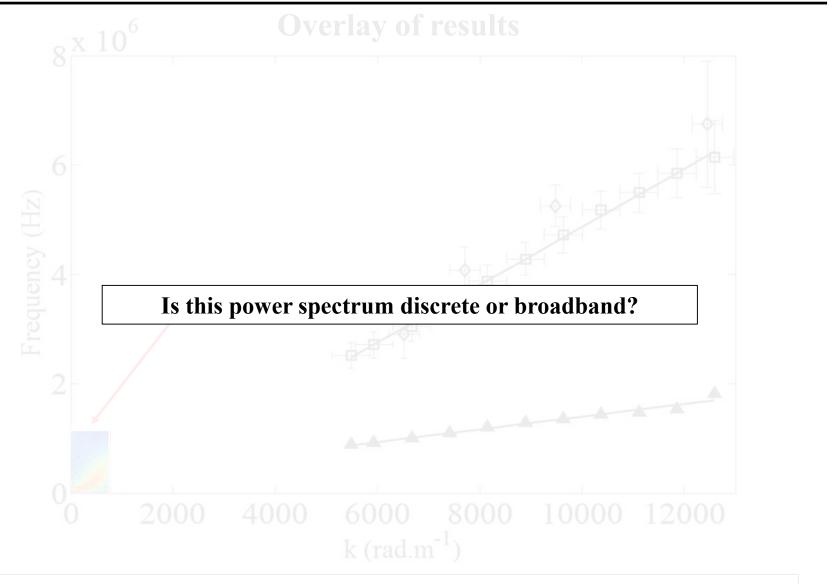






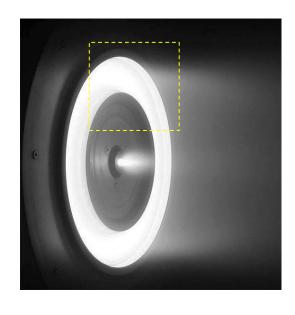
Comparison is not one to one but does illustrate both follow similar linear trends

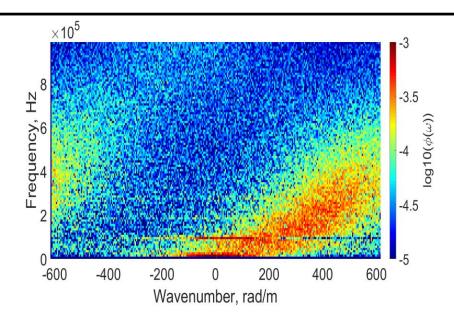




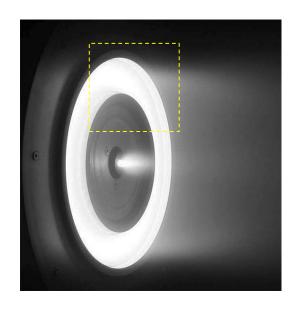
Comparison is not one to one but does illustrate both fall on similar linear trends

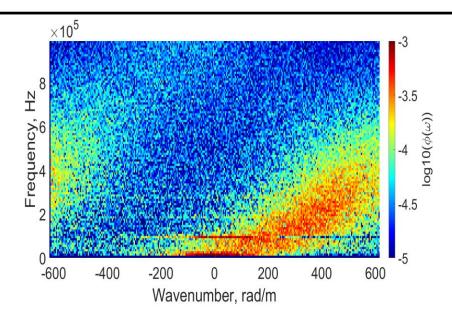


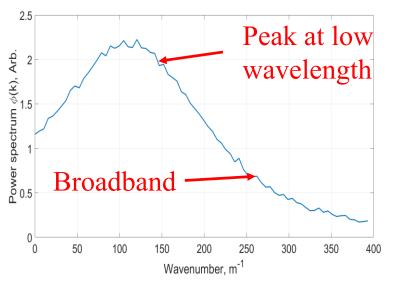




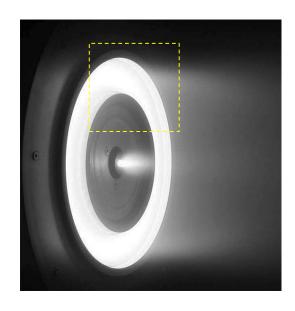






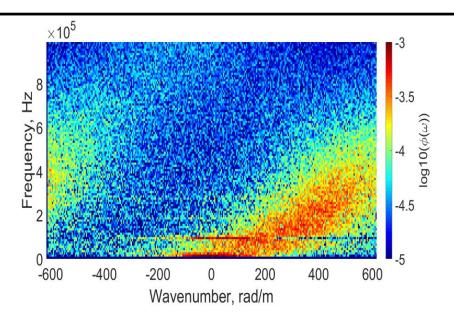


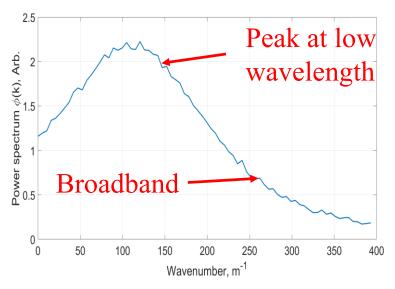




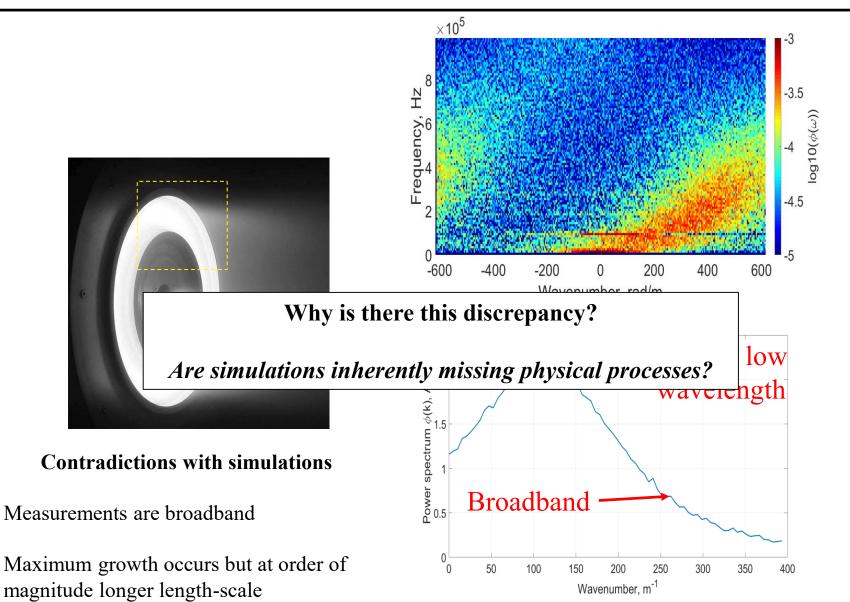
Contradictions with simulations

- Measurements are broadband
- Maximum growth occurs but at order of magnitude longer length-scale



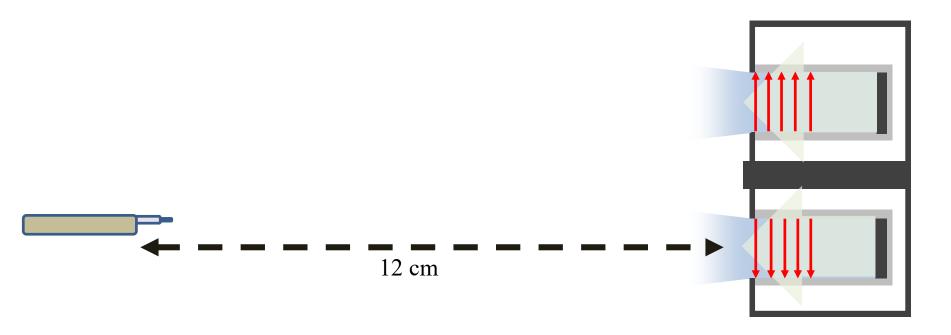






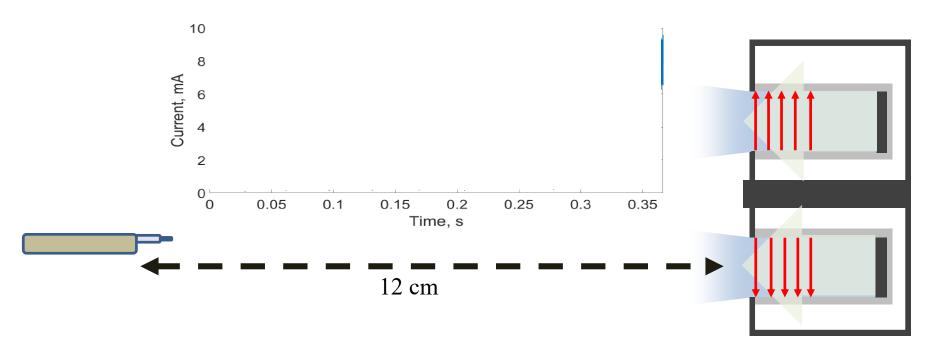


- Record continuously as probe is injected at high speed
- Chop waveform into position bins based on injection trajectory



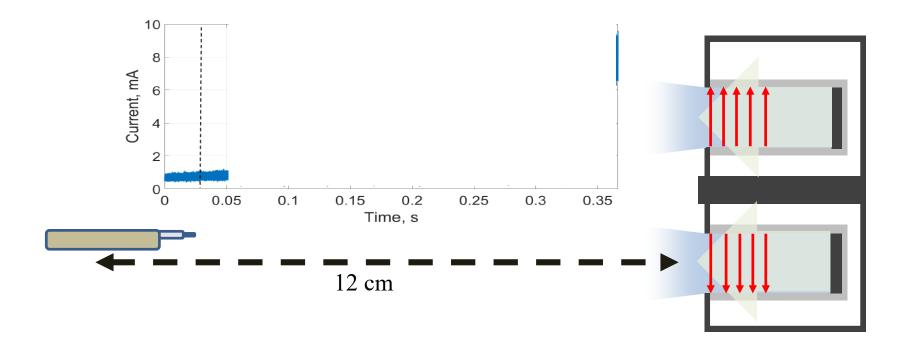


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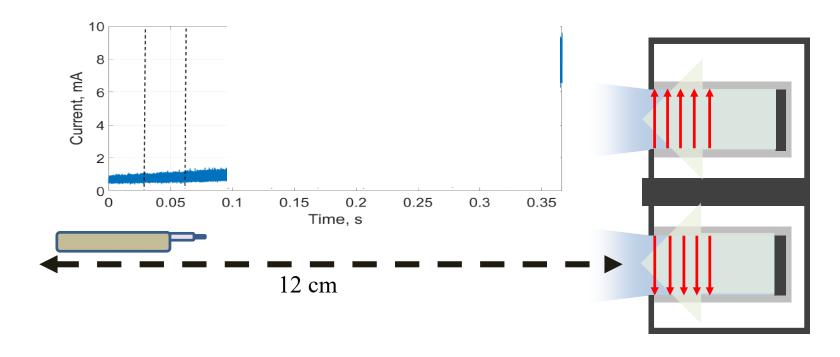


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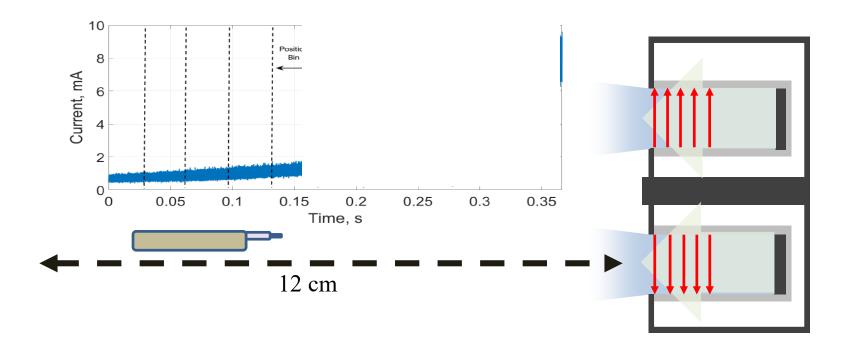


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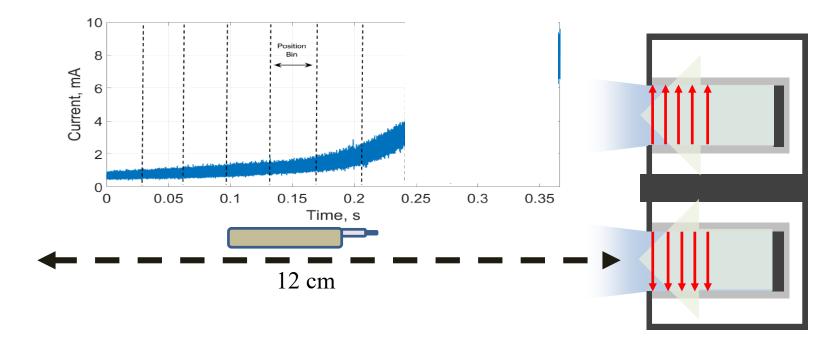


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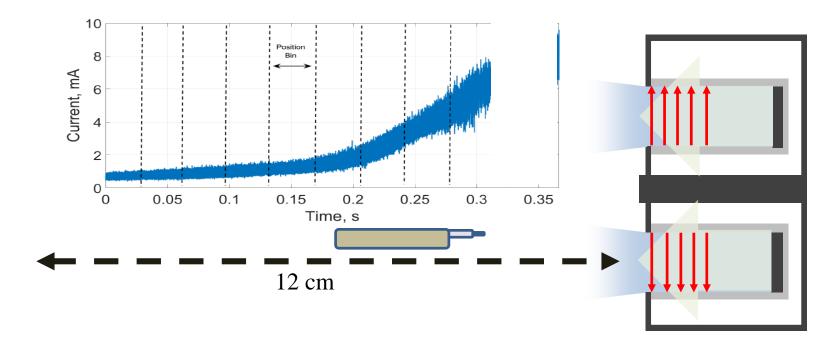


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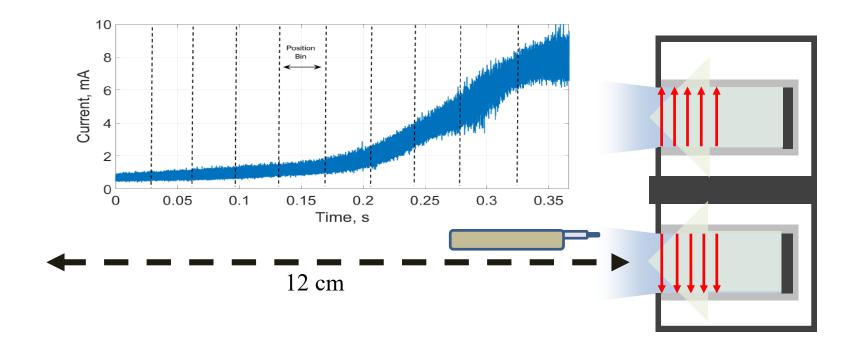


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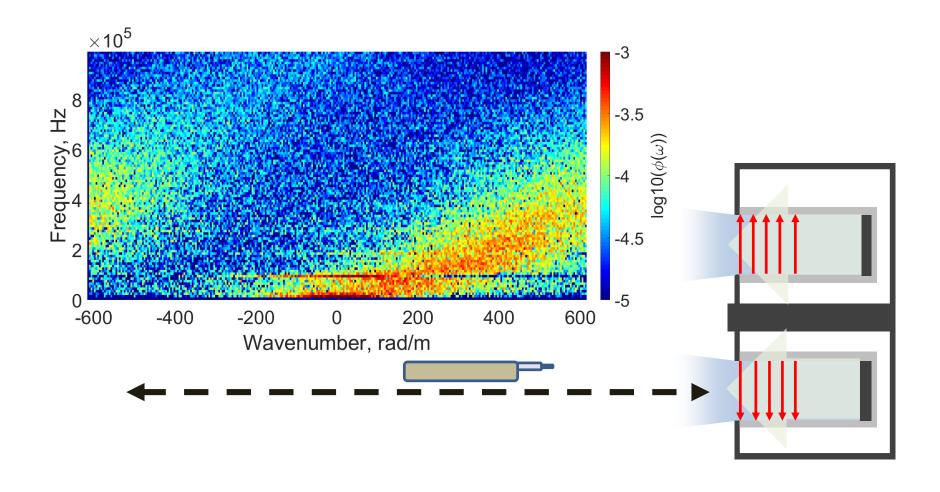




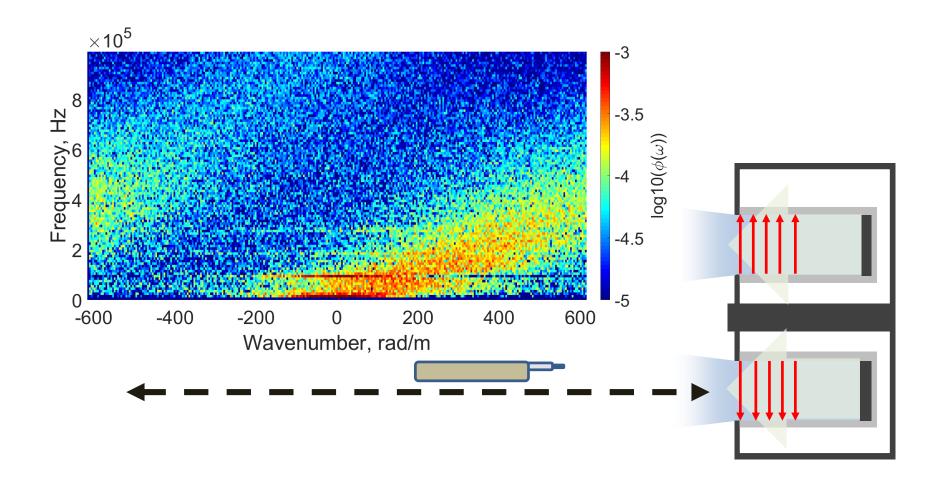
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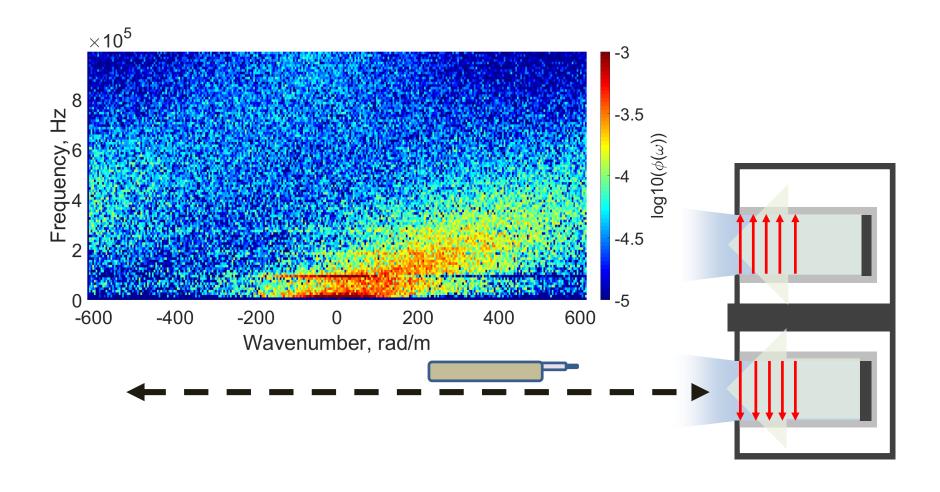




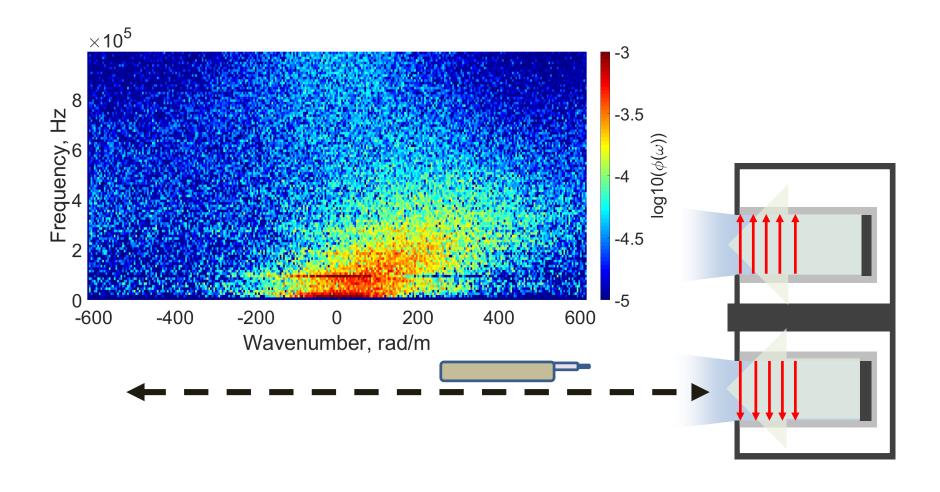




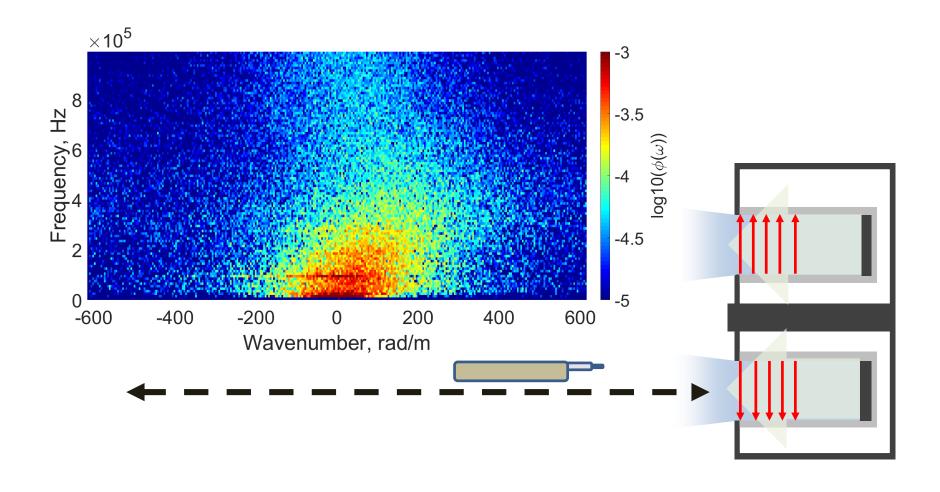




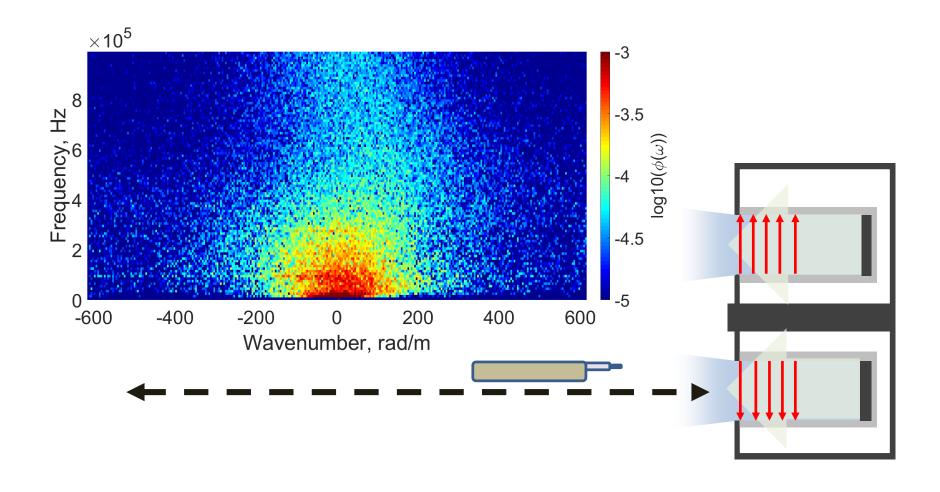




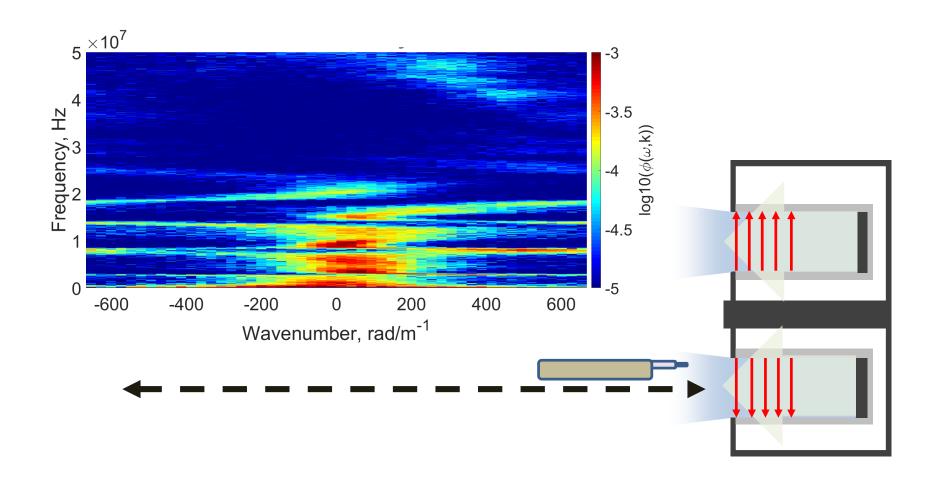








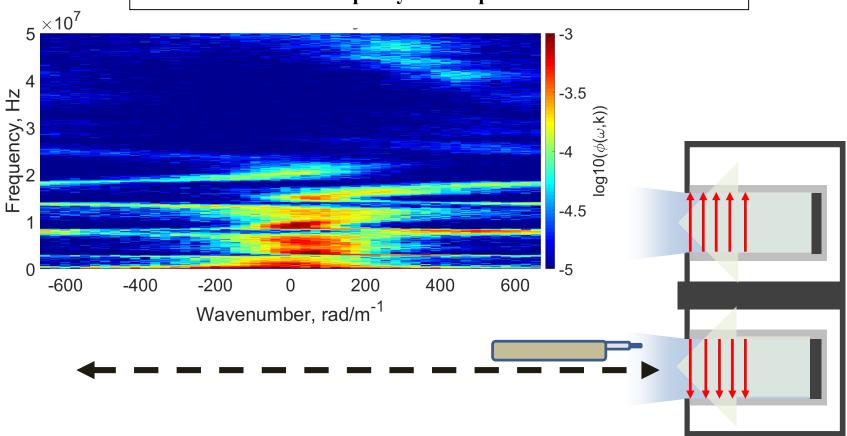




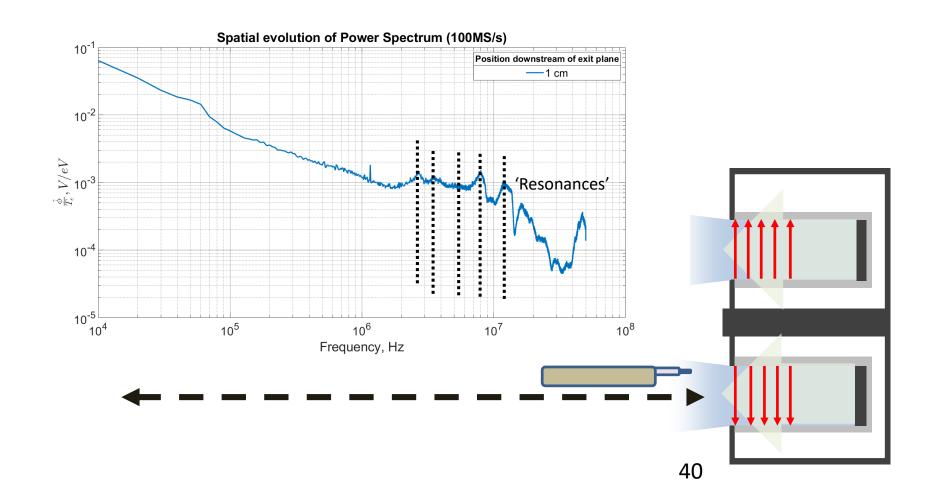


The dispersion relation fundamentally changes nature showing discrete structure upstream

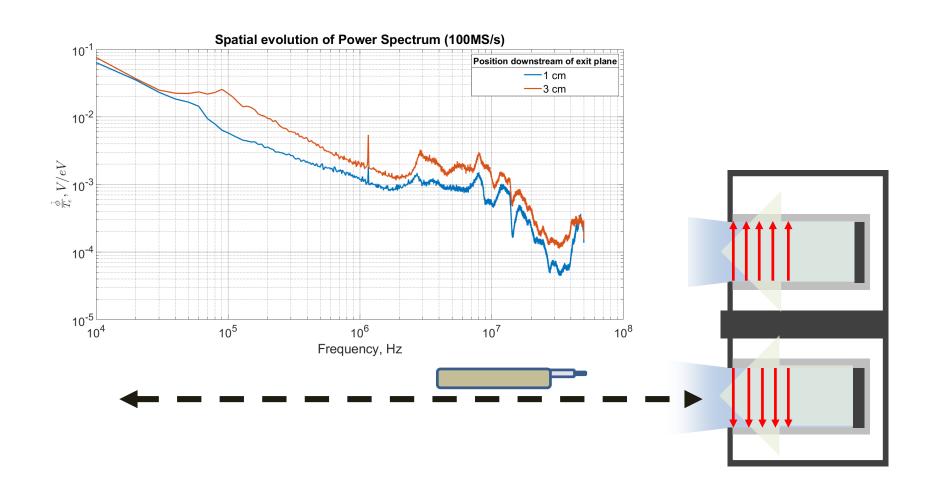
Could discrete structure transitioning into broadband explain discrepancy with experiments?



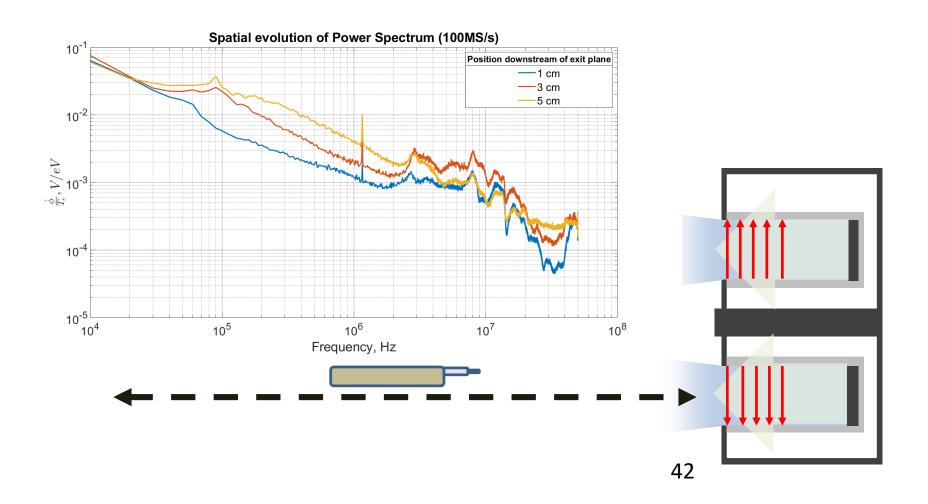




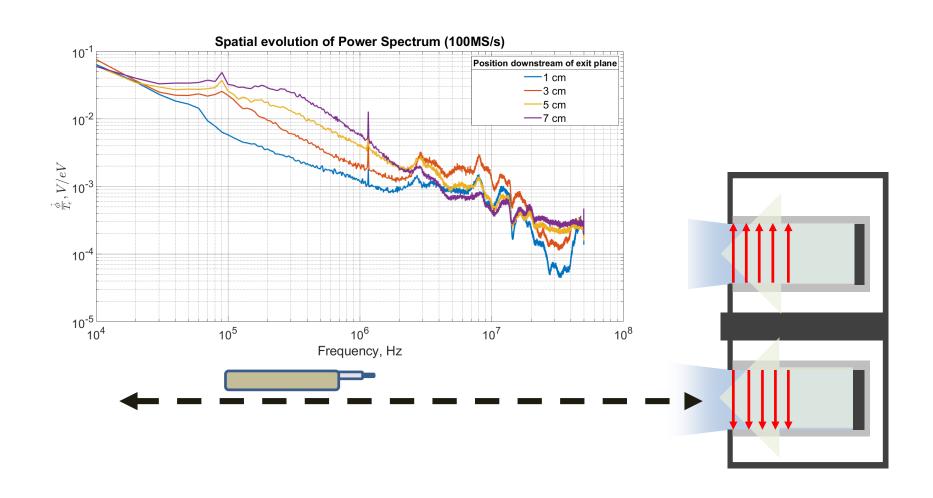




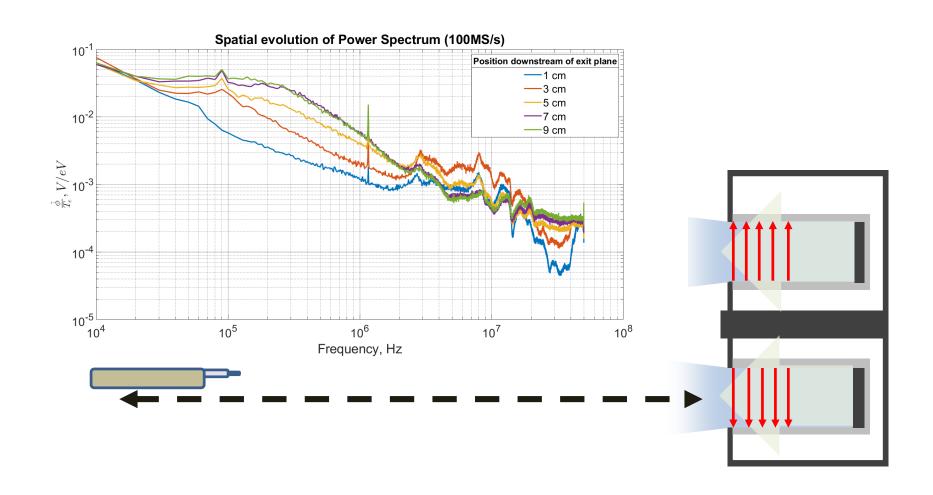




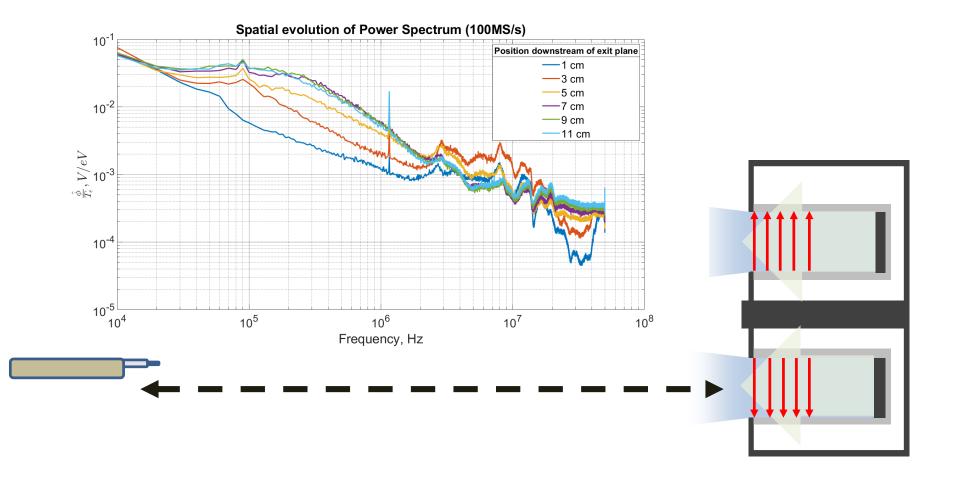




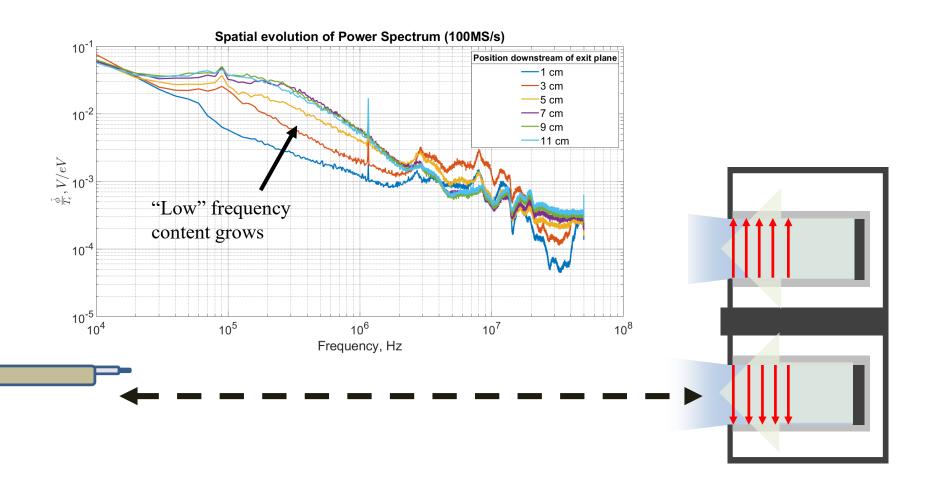








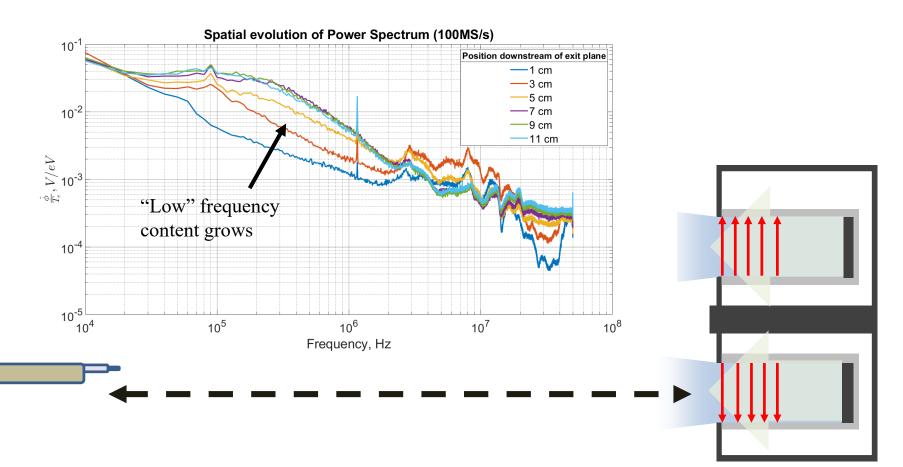






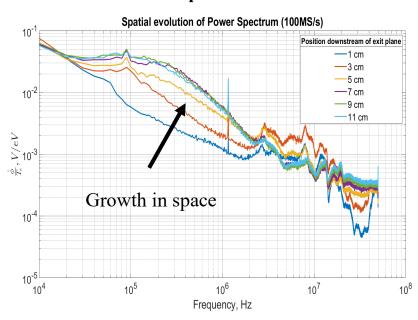
Hypothesis: experiments to date have been performed too far downstream of acceleration zone to image formation of ECDI where discrete structure is present

How could this transition occur?



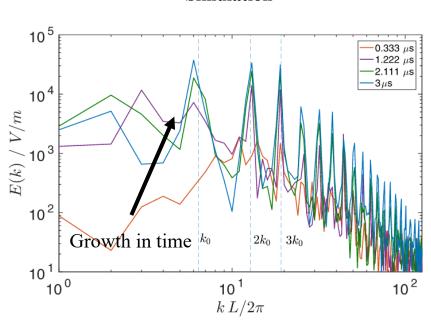


Experiment



Z. Brown and B. Jorns (submitted) 2018.

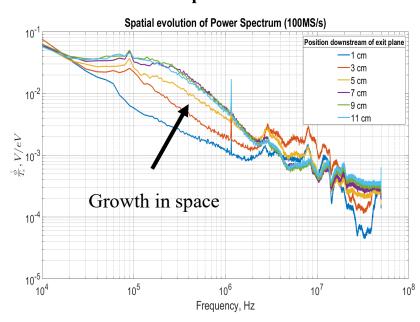
Simulation



S. Janhunen et al., Physics of Plasmas, 011608 (2018)

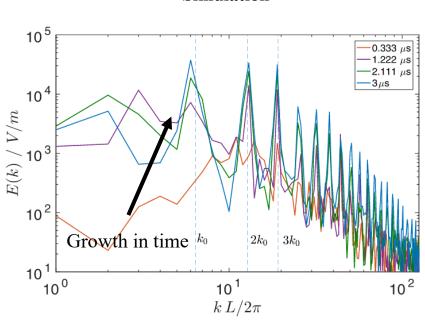


Experiment



Z. Brown and B. Jorns (submitted) 2018.

Simulation

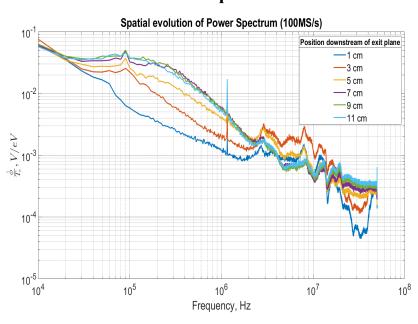


S. Janhunen et al., *Physics of Plasmas*, 011608 (2018)

Could there be a non-linear inverse energy transfer from short to long wavelength?



Power spectra



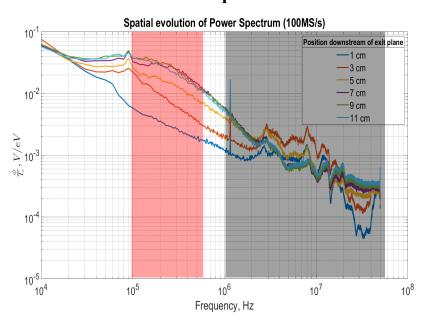
Z. Brown and B. Jorns (submitted) 2018.

Increase in low wavelength content energy (acoustic modes) increases with position as energy in resonant modes decreases

Efforts underway to quantify and understand this transfer



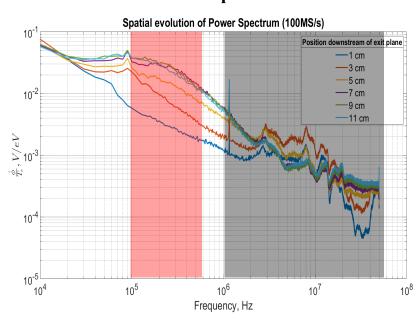
Power spectra



Z. Brown and B. Jorns (submitted) 2018.

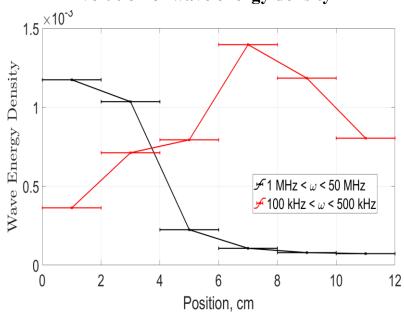


Power spectra



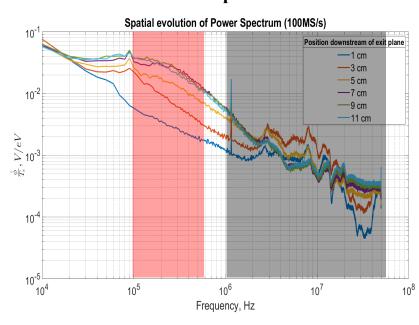
Z. Brown and B. Jorns (submitted) 2018.

Evolution of wave energy density



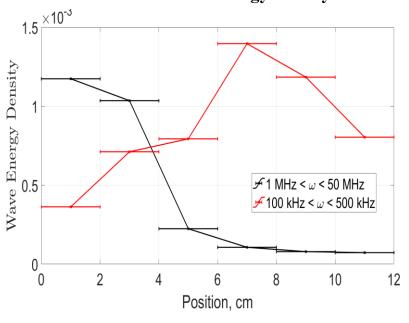


Power spectra



Z. Brown and B. Jorns (submitted) 2018.

Evolution of wave energy density

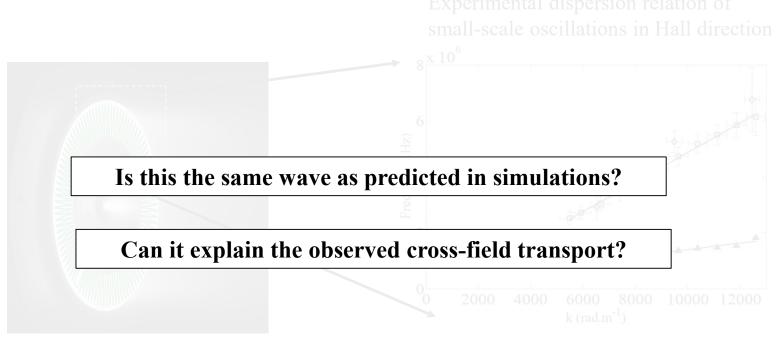


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Experimental evidence of microturbulence



S. Tsikata, N. Lemoine, V. Pisarev, and D. Grésillon, *Physics of Plasmas*, Vol. 16, No. 3, 2009

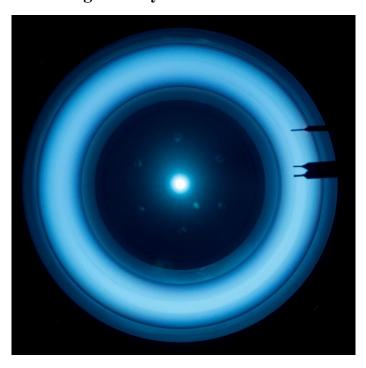
- Wavelengths < 1 mm
- Dispersion is acoustic-like

ECDI in the acoustic-like limit

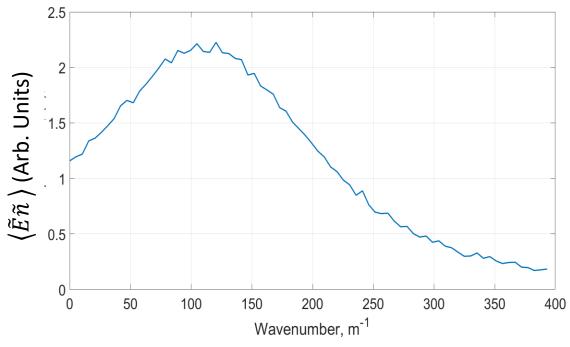
Modes are incoherent



9kW Magnetically Shielded Hall thruster



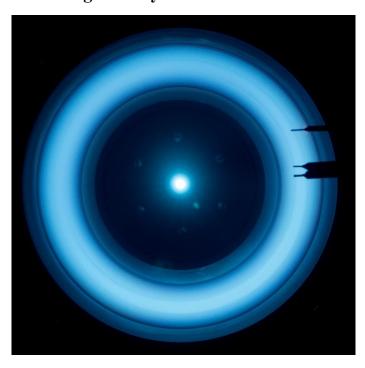
Power spectrum of ECDI-like oscillations



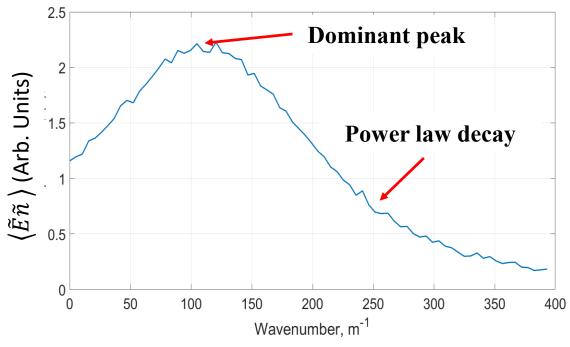
Z. Brown and B. Jorns, "Dispersion relation measurements of plasma modes in the near-field plume of a 9-kW magnetically shielded thruster," IEPC-2017-387



9kW Magnetically Shielded Hall thruster



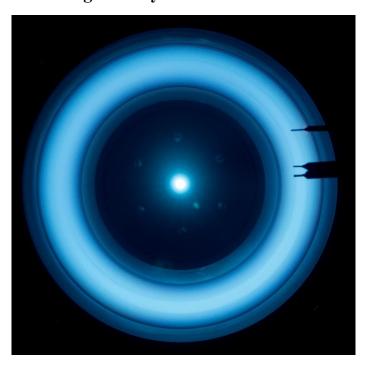
Power spectrum of ECDI-like oscillations



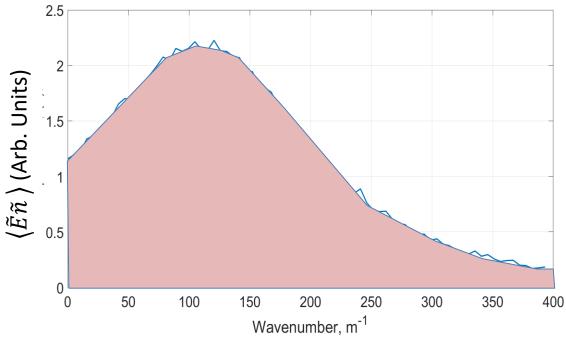
Z. Brown and B. Jorns, "Dispersion relation measurements of plasma modes in the near-field plume of a 9-kW magnetically shielded thruster," IEPC-2017-387



9kW Magnetically Shielded Hall thruster



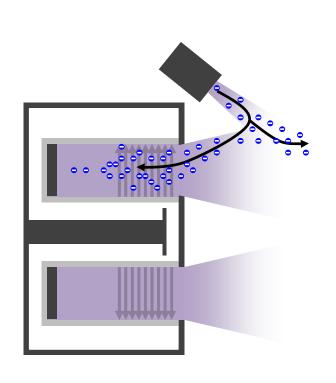
Power spectrum of ECDI-like oscillations



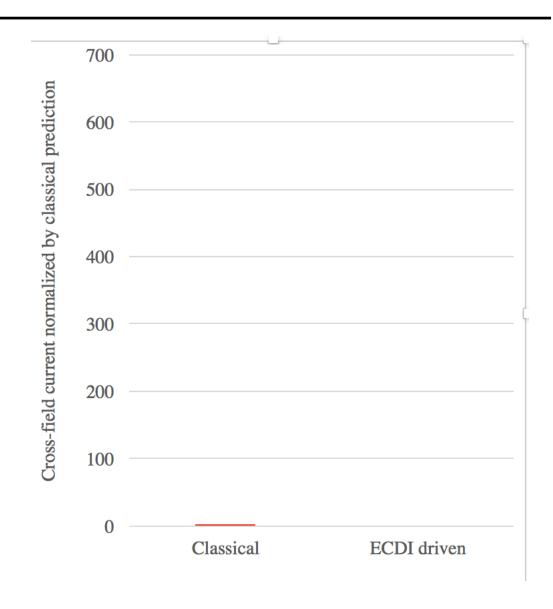
Z. Brown and B. Jorns, "Dispersion relation measurements of plasma modes in the near-field plume of a 9-kW magnetically shielded thruster," IEPC-2017-387

$$I_e \propto \frac{q}{B} \langle \tilde{E} \tilde{n} \rangle$$

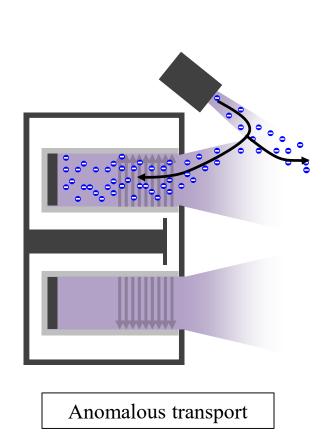


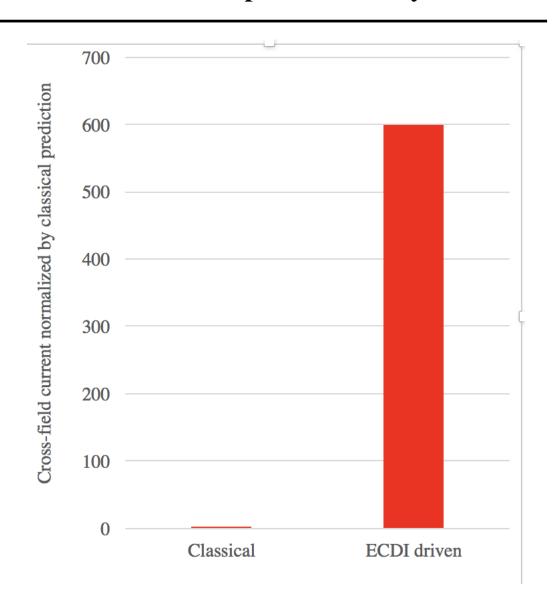


Classical transport from particle collisions

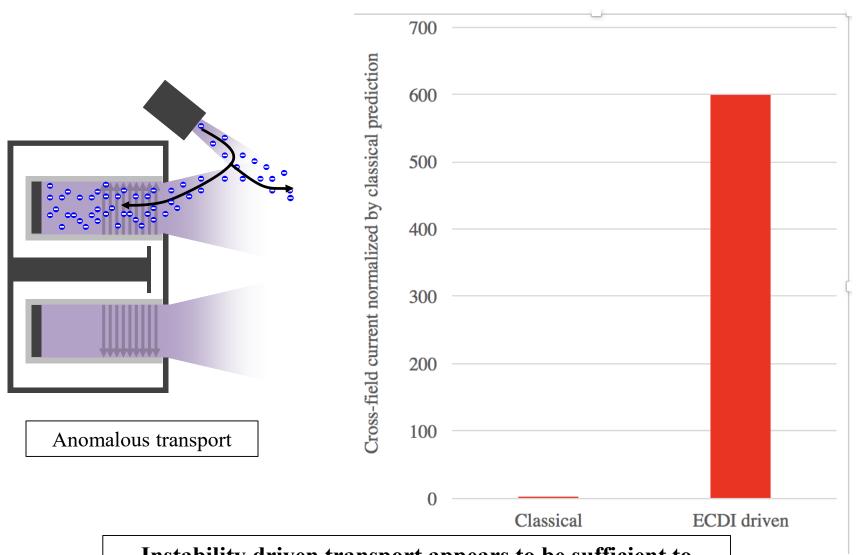








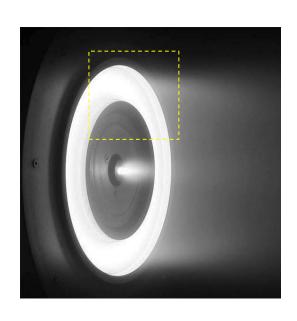




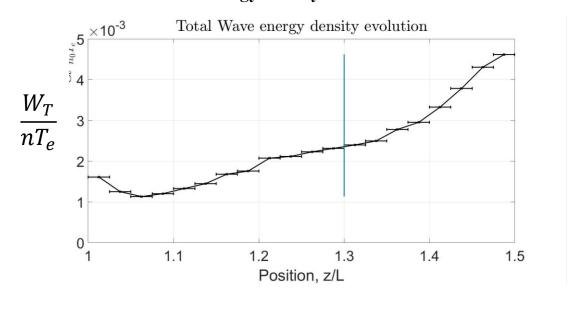
Instability driven transport appears to be sufficient to explain electron dynamics in plume (with caveats)



Other experimental insights into ECDI



ECDI energy density in thruster channel



$$v_{AN} = \alpha \gamma_e \frac{W_T}{nT_e} \approx \alpha \gamma_e$$

Critical information for fluid-closure



Summary

- A number of simulations suggest that ECDI exists and should exhibit distinct, measureable features
 - Discrete
 - Maximum growth at small wavelength
- Experimental measurements to date have not supported this
- Probing has shown that discrepancy may in part be due to fact ECDI spectrum transitions from upstream discrete nature to downstream broadband spectrum
- The nature of this transition is not known but is currently being studied
- Other experimental insights are guiding investigations into closure, e.g. the observation modes may be saturated



Follow on experimental questions

- How does the energy cascade occur and can we measure it?
- Can we measure shorter wavelengths inside the channel with necessary spatial resolution?
- Can we use experiments to guide closure for simulation results?
- Can kinetic codes ever capture all the nuanced effects (like nonlinear energy cascade in 3D) likely occuring in this geometry?
- What other experimental measurements can we make to characterize predictions related to ECDI:
 - Ion azimuthal drift (on-going at UM)
 - Direct measurements of electron transport (on-going at UM)